1	Recording Fidelity of Relative Paleointensity Characteristics in the North Pacific										
2	Ocean										
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15											
16	Key Points:										
17	• We use an age model based on tephra correlation and one radiocarbon date to assess										
18	the fidelity of relative paleointensity variations.										
19	• Detrital vortex state and biogenic single domain magnetite are the dominant										
20	remanence carriers, and they introduced <0.5-kyr age error.										
21	• Assessment of component-specific magnetic properties should be considered to										
22	ensure the reliability of relative paleointensity variations.										
23											

### 24 Abstract

North Pacific Ocean sediments are important archives of geomagnetic field and 25 paleoenvironmental evolutions, yet late Quaternary sedimentary sequences in the North Pacific 26 27 Ocean are less explored because of generally low sedimentation rates and the challenge of dating sediments deposited below the calcite compensation depth. Core NP02 sediments from the North 28 Pacific Ocean contain three visible tephra layers, which are identified as the To-Of, Spfa-1, and 29 Kt-3 tephra, respectively. An age model for core NP02 is established based on tephra correlation 30 and one radiocarbon date, which gives a mean sedimentation rate of ~11.6 cm/kyr. Magnetic 31 analyses suggest the dominance of two remanence carriers; detrital vortex state and biogenic single 32 domain magnetite. Variable depth offsets between the relative paleointensity (RPI) signals carried 33 34 by the two remanence carriers adds a negligible age error (<0.5 kyr) to the core NP02 RPI record. 35 Pronounced RPI amplitude variations are observed, and the Rockall and Laschamp excursions are 36 identified in the core NP02 RPI record based on the age model. The fidelity of the core NP02 RPI 37 record is further verified by comparison with other RPI stacks and records, and we demonstrate 38 that the RPI variations can assist in chronology refinement.

39

## 40 Plain language Summary

The North Pacific Ocean is too deep to preserve materials effectively for the oxygen isotope chronology, and sediments deposit slowly in this region. Therefore, dating is a big problem for late Quaternary sedimentary cores. In this paper, core NP02 sediments from the North Pacific Ocean were studied. The To-Of, Spfa-1, and Kt-3 tephra layers were identified in core NP02 sediments based on the geochemical features. An age model for this core is reconstructed with the aid of tephra ages and one radiocarbon date. Magnetic analyses suggest that both small biogenic 47 and relatively large detrital magnetite record geomagnetic information, and the two groups of 48 magnetite deposit at about the same time. Based on the age model, core NP02 RPI record preserves 49 the Rockall and Laschamp excursions, and the RPI pattern is consistent with other RPI stacks and 50 records, which verified core NP02 RPI fidelity. Therefore, we demonstrate that RPI variations is 51 a promising dating tool for North Pacific Ocean sediments.

52

## 53 **1 Introduction**

As the largest ocean in the world, the Pacific Ocean preserves abundant information about 54 geomagnetic field and paleoenvironmental evolutions (e.g., Channell & Lanci, 2014; Doell & Cox, 55 1971; Dunlea et al., 2017; Jacobel et al., 2017; Lanci et al., 2004; McKinley et al., 2019; Rea et 56 al., 1998; Roberts et al., 1997; Yamazaki & Yamamoto, 2018). However, sedimentation rates in 57 the North Pacific Ocean are generally low (e.g., Korff et al., 2016; Kyte et al., 1993; Ninkovich et 58 59 al., 1966; Pettke et al., 2000; Prince et al., 1980; Rea et al., 1998), and sediments in this region are usually deposited below the calcite compensation depth (CCD), which impedes development of 60 radiocarbon dating and continuous oxygen-isotope stratigraphy to construct precise age models to 61 assist understanding of late Quaternary geological events. Therefore, a lot of short sedimentary 62 cores in the North Pacific Ocean have been less investigated, and there is an impetus to provide a 63 robust and practical way to date these sediments, and then promote the relevant studies. 64

Relative paleointensity (RPI) variations of Earth's magnetic field have been used as a
chronological tool in marine sediments (e.g., Channell et al., 2009; Channell, Hodell, et al., 2016;
Guyodo & Valet, 1996, 1999; Valet et al., 2005; Yamamoto et al., 2007). This makes RPI
variations valuable in environments like the deep North Pacific Ocean (Roberts et al., 1997; Shin

et al., 2019; Yamamoto et al., 2007; Yamazaki, 1999). The main advantage of the RPI approach 69 is that dipole geomagnetic variations are recorded synchronously around the world and represent 70 a geophysical measurement that is independent of seawater chemistry and that can be used to 71 develop independent age constraints (Roberts et al., 2013). Paleointensity-assisted chronology 72 studies are based on the assumption that sedimentary magnetic records contain information about 73 the magnetic field intensity at the time of sediment deposition (Tauxe, 1993). To ensure that RPI 74 records are not contaminated by other sedimentary magnetic signals, magnetite should be the 75 dominant remanence carrier, the concentration of magnetite should vary down-core by less than 76 one order of magnitude, and the grain size of magnetite should be in the 1-15 µm range (King et 77 al., 1983; Tauxe, 1993). However, these requirements are based on the assumption that sediments 78 79 contain a single magnetic mineral component rather than a mixture of several components. Recent studies suggest that magnetite with different grain-sizes responds differently to the geomagnetic 80 field during paleomagnetic signal acquisition in sediments (Chen et al., 2017; Ouyang et al., 2014; 81 82 Roberts et al., 2012, 2013; Yamazaki, Yamamoto, et al., 2013). Biogenic and terrestrial magnetite co-occur pervasively in North Pacific Ocean sediments (Yamazaki & Ioka, 1997; Yamazaki & 83 84 Shimono, 2013; Q. Zhang et al., 2018). Given that biogenic and terrestrial magnetite have different 85 grain-size distributions, potential complexities in magnetic recording need to be considered, 86 including the possible presence of different remanence acquisition mechanisms with different lock-in depths. Therefore, detailed assessment of magnetic grain-size characteristics should be 87 made to test the fidelity of sedimentary RPI records. In this study, we use an independent age 88 89 model for North Pacific Ocean core NP02 to assess whether RPI correlations based on magnetite components with different grain-size distributions can be used to refine the core chronology. 90

### 91 2 Materials and Methods

92 **2.1 Core description and sampling** 

Core NP02 is a 4.35 m gravity piston core from 40.48° N, 150.1° E. It was recovered from the North Pacific Ocean abyssal plain at a water depth of 5,177 m. This location lies below the CCD, so few foraminifera are preserved. The core site is downwind of Asian dust source areas and many volcanoes in the Japanese archipelago and Kamchatka Peninsula (Weber et al., 1996) (Figure 1). Therefore, the studied sediments are composed mainly of clay with occasionally intercalated tephra layers. Cubic samples (8 cm<sup>3</sup>) were taken continuously from the working half of core NP02 for paleomagnetic and rock magnetic analyses.





## 105 **2.2 Methods**

#### 106 2.2.1 Magnetic measurements

Magnetic remanence measurements were made using a 2-G Enterprises Model 760R 107 cryogenic magnetometer with an in-line alternating field (AF) demagnetizer. The natural remanent 108 109 magnetization (NRM) was measured and stepwise demagnetized to a peak field of 100 mT, in 110 logarithmically increasing field steps of 0.1-7.7 mT. Characteristic remanent magnetization (ChRM) directions were determined by principal component analysis (PCA) with successive 111 112 demagnetization points that define vectors directed toward the origin of demagnetization diagrams 113 (Kirschvink, 1980). Core NP02 is not oriented, therefore, the mean ChRM declination was calculated by making a constant azimuthal rotation to all samples to yield an adjusted mean 114 115 declination of 0°. An anhysteretic remanent magnetization (ARM) was imparted in a 100 mT peak field with a 0.05 mT direct current (DC) bias field, and was then demagnetized with applied fields 116 of 10, 20, 30, 40, 50, 60, 70, 80, 90, and 100 mT. Isothermal magnetizations (IRMs) were imparted 117 in a DC field of 1 T, which is defined here as the saturation IRM (SIRM). 118

119 Magnetic susceptibility ( $\chi$ , mass specific) was measured at a frequency of 976 Hz. 120 Temperature-dependent magnetic susceptibility ( $\chi$ -*T*) curves for representative samples were 121 measured in an argon atmosphere from room temperature to 700 °C.  $\chi$  and  $\chi$ -*T* curves were 122 measured using an AGICO Kappabridge MFK 1-FA system equipped with a CS-3 furnace in an 123 argon atmosphere.

IRM acquisition curves and first-order reversal curves (FORCs) (Pike et al., 1999; Roberts
 et al., 2000) were measured to a maximum field of 1 T using a MicroMag Model 3900 vibrating
 sample magnetometer. FORC diagrams were processed with the FORCinel software v3.05

- (Harrison & Feinberg, 2008), and IRM acquisition curves were decomposed following Robertsonand France (1994).
- 129

## 2.2.2 Electron microscope analyses

Magnetic extracts from selected samples were carbon-coated or loaded onto a carboncoated copper grid for electron microscope analyses. Scan electron microscope (SEM) and transmission electron microscope (TEM)/scanning transmission electron microscope (STEM) observations were made with a PHENOM XL desktop SEM at 15 kV and a F200S S/TEM at 200 kV, respectively. Energy-dispersive X-ray spectroscopy (EDS) analysis was performed to determine the chemical compositions of selected areas semi-quantitatively.

136  $2.2.3^{14}C$  dating

<sup>14</sup>C dates were obtained using accelerator mass spectrometry on planktonic foraminifera (*N. pachyderma* sinistral). A probability density function of calibrated ages was calculated using <sup>14</sup>C determinations and the 1 $\sigma$  range from the Marine13 calibration curve (Reimer et al., 2013) <sup>14</sup>W with a reservoir age of  $\Delta R = 400\pm50$  years based on Yoneda et al. (2007). This calibrated age <sup>14</sup>distribution was analyzed subsequently using the Bayesian highest posterior density method to <sup>14</sup>produce a calibrated age range determination following Lougheed and Obrochta (2016).

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## 2.2.4 Volcanic glass shard analysis

Samples from the identified tephra layers were treated with  $H_2O_2$ , and then volcanic glass shards were selected, mounted, ground, and polished for electron probe microanalysis. Major element compositions were measured with a JEOL JXA-8100 electron microprobe at 15 kV accelerating voltage with 10  $\mu$ m beam diameter and 6 nA beam current. Na contents were determined first to reduce the impact of its mobilization. Secondary standard glass, ATHO-G and

- 149 StHS6/80-G, were used to check the precession and accuracy of the data (Jochum et al., 2006). For
- 150 comparison, all data were normalized to an anhydrous basis (i.e., 100% total oxides).
- 151

152 **3 Results** 

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## 3.1 Paleomagnetic stability

The NRM is demagnetized almost completely at 100 mT (Figure 2a), which suggests that 154 the NRM is carried mainly by low-coercivity magnetic minerals. A secondary remanent 155 magnetization is present in some samples and can be removed by AF demagnetization to 10 mT. 156 ChRMs can be identified as components directed toward the origin of demagnetization diagrams 157 up to 100 mT. Maximum angular deviation (MAD) values were calculated by PCA (Kirschvink, 158 1980) and are generally  $< 5^{\circ}$  (Figure 2b), which suggests that the ChRMs are stable and well-159 defined. ChRM inclinations vary about an average value of 52°, which is slightly shallower than 160 the expected geocentric axial dipole value for this site ( $\pm 59.6^{\circ}$ ). Shallower inclinations may be 161 162 associated with acquisition of a depositional remanent magnetization or sediment compaction (Anson & Kodama, 1987; Arason & Levi, 1990; Roberts et al., 2013; Tauxe & Kent, 2004). Three 163 successive reversed polarity inclination values are identified at depths of 39 to 45 cm, which are 164 dinterpreted to represent a geomagnetic excursion. 165



Figure 2. (a) Representative orthogonal projections of AF demagnetization data for sediments from core
NP02. Solid (open) circles denote projections of the magnetization vector onto horizontal (vertical) planes.
(b) ChRM inclinations, declinations, and MAD values. Dashed lines represent the expected inclination for
a geocentric axial dipole field at the site latitude and inclination = 0°, respectively.

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# 3.2 Magnetic mineralogy

Bulk rock magnetic results are shown in Figure 3.  $\chi$ -*T* heating curves increase steadily from room temperature to ~300-400 °C (Figure 3a), which indicates gradual unblocking of single domain (SD) particles and/or heating-induced ferrimagnetic mineral neoformation.  $\chi$ -*T* curves for tephra-containing samples (i.e., at depths of 141 and 393 cm in Figure 3a) have larger  $\chi$  values than other curves during both heating and cooling. The heating curve for one tephra-containing sample (141 cm) reaches a near-zero value at ~660 °C, which suggests the presence of hematite. Otherwise, all heating curves decrease to near-zero values at ~580 °C, which corresponds to the

Curie temperature of magnetite (Dunlop & Özdemir, 1997). Down-core variations of 178 179 concentration-dependent magnetic parameters ( $\chi$ , ARM, and SIRM) are shown in Figure 3b. Tephra-containing samples are distinguished by anomalously high  $\chi$ , SIRM, and ARM values, and 180 were removed from our paleomagnetic record. Once these tephra-dominated intervals are removed, 181  $\chi$ , ARM, and SIRM values vary within a factor of 5.5, 5.5, and 6.8, respectively, which suggests 182 that the magnetite concentration is relatively homogeneous (i.e. variation of less than one order of 183 184 magnitude) throughout core NP02. ARM/SIRM is used widely as a magnetic grain-size proxy. However, ARM acquisition can be affected strongly by the magnetic mineral concentration, with 185 higher magnetite concentrations reducing ARM acquisition efficiency (King et al., 1983; Sugiura, 186 187 1979). If SIRM is used to represent the magnetic mineral concentration, a higher SIRM should produce lower ARM/SIRM values. A bi-plot of ARM/SIRM and SIRM for core NP02 does not 188 have such a trend (Figure 3c), which indicates that the magnetite concentration has no significant 189 effect on ARM acquisition efficiency. Therefore, ARM/SIRM can be employed here as a magnetic 190 191 grain-size proxy. A bi-plot of ARM and SIRM has a linear trend with little scatter (Figure 3d), which suggests that changes in bulk magnetite grain size are relatively minor. 192



Figure 3. Rock magnetic results from core NP02. (a) Temperature dependence of magnetic susceptibility ( $\chi$ -T) curves; left (right) are heating (cooling) curves (the depths of the representative samples are indicate in the legend); (b) down-core concentration parameter variations for  $\chi$ , ARM, and SIRM; major ash layers are indicated by grey shaded areas; (c) bi-plot of SIRM and ARM/SIRM; and (d) bi-plot of SIRM and ARM. Brown circles in (c) and (d) represent volcanic ash-containing samples.

198 Although the bulk magnetic properties are relatively homogenous, component-specific analyses suggest the presence of mixed magnetic signatures in core NP02. FORC diagrams contain 199 a ridge-like distribution along the  $B_c$  axis with little vertical spreading, along with an 200 asymmetrically distributed component that diverges toward the  $B_u$  axis (Figures 4a-4f). The former 201 component together with a negative contribution along the  $B_u$  axis (Figures 4a-4d) is due to SD 202 particles with uniaxial anisotropy (Muxworthy et al., 2004; Newell, 2005), which is often 203 associated with non-interacting biogenic SD particles (Egli et al., 2010), while the latter 204 component is typical of detrital particles in the magnetic vortex state (Lascu et al., 2018; 205 206 Muxworthy & Dunlop, 2002; Roberts et al., 2017). On the B<sub>c</sub> axis, contours can extend to 110 mT,



Figure 4. Magnetic properties of typical samples from core NP02. (a-f) FORC diagrams with VARIFORC (Egli, 2013) smoothing parameters of  $S_{c,0} = 6$ ,  $S_{c,1} = 9$ ,  $S_{b,0} = 6$ ,  $S_{b,1} = 9$ ; (g-l) IRM acquisition curve decomposition. Blue, red, purple, grey, and green curves are for components 1, 2, 3, 4, and the sum of components, respectively; and (m) down-core variations of IRM for components 1 and 2 derived from IRM acquisition curve decomposition.

212 which indicates the presence of high-coercivity magnetic particles (e.g., Ahmadzadeh et al., 2018; Gai et al., 2020). Some FORC diagrams also contain a secondary peak near the origin (Figures 4a, 213 4b, and 4d), which indicates viscous behaviour associated with particles near the 214 215 superparamagnetic/stable SD (SP/SSD) threshold size (Pike et al., 2001; Roberts et al., 2000). Magnetic particles in different grain sizes were verified by the electron microscopic observations, 216 which can be interpreted as detrital vortex and biogenic SD magnetic particles based on their 217 218 characteristic morphologies and grain sizes (Figure 5). Magnetite seen in SEM images are submicrons to several microns in length, and their shapes are irregular, which are typical of detrital 219 origin (e.g., Ao et al., 2012) (Figure 5e-5g). In comparison, biogenic magnetite seen in TEM 220 images are several tens of nanometers, which are within the SD range. They are evidenced by the 221 222 regular elongated or equant morphologies (Figure 5a-5d), and by the intact magnetosome chain structure (Figure 5b). Apart from the biogenic magnetite, SD magnetite in TEM images are also 223 224 occurred as inclusions within detrital silicate host; these magnetic inclusions are clustered (Figure 225 5h, 5j) or dentritic textures (Figure 5i). For IRM acquisition curves, samples were not saturated at 226 the field of 300 mT, which suggests the presence of high-coercivity magnetic minerals (Figure 227 S1); the asymmetric gradient (Figures 4g-4l) suggests a mixing of multiple coercivity/mineral phases that is consistent with indications from the FORC diagrams. Four components are identified 228 229 through IRM acquisition curve decomposition of the selected 51 samples. Component 1 ( $B_{1/2}$  = 24-33 mT, DP = 0.20-0.22) and component 2 ( $B_{1/2}$  = 54-69 mT, DP = 0.19-0.21) contribute more 230 than 75% of the IRM. Considering that magnetic inclusions in silicates contribute little to 231 remanence in oxic pelagic Pacific sediements (Usui et al., 2018), components 1 and 2 are 232 associated with the detrital vortex state and biogenic SD components indicated in FORC diagrams, 233 respectively. Component 3 ( $B_{1/2} = 148-182 \text{ mT}$ , DP = 0.20-0.23) is similar to the low- and medium-234

coercivity components identified in Chinese loess-paleosol sequences (Hu et al., 2013; Spassov et al., 2003; R. Zhang et al., 2016), which is interpreted to represent detrital maghemite and/or pedogenic hematite. Component 4 ( $B_{1/2} < 10 \text{ mT}$ , DP = 0.26-0.27) may represent pedogenic magnetic particles just above the SP/SSD threshold size, which correspond to the component indicated by the secondary peak at the origin of FORC diagrams. Downcore, IRMs of components 1 and 2 vary within a factor of 4.2 and 5.8, respectively (Figure 4m).



Figure 5. Electron microscope images of particles in magnetic extracts from core NP02. (a-d) TEM images for biogenic magnetite; (e-g) SEM images for detrital magnetite; and (h-j) STEM images for magnetic inclusions within silicates. Red (c), blue (g), green and orange (i) crosses indicate selected areas for EDS analysis, which is presented in (k). EDS analysis indicates that magnetic inclusions and their hosts are magnetite (orange spectrum) and silicates (green spectrum), respectively. The Cu peaks originate from the TEM copper grid. Magnetite particles identified by "Mt" have been verified via EDS.

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## 3.3 Normalized remanence records

248 RPI values can be estimated by normalizing the NRM by a laboratory-induced 249 magnetization, such as IRM or ARM, to compensate for magnetic mineral concentration changes. 250 Selection of a suitable RPI normalizer depends largely on the grain size distribution of the 251 remanence carrier. ARM is mainly carried by finer magnetic particles, while IRM is also carried 252 by coarser MD particles. Core NP02 samples are dominated by fine (SD and vortex state) magnetic 253 particles, so ARM is chosen as the normalizer.

254 Our paleomagnetic and rock magnetic results satisfy conventional criteria for RPI studies (King et al., 1983; Tauxe, 1993), and further indicate that two components (hypothesized to be 255 detrital and biogenic magnetite) with different magnetic grain size distributions dominate the 256 magnetic mineral assemblage. The coercivities of the detrital and biogenic components do not 257 overlap strongly over the 10-40 and 50-100 mT windows (blue and pink bands, respectively, in 258 Figures 4g-41). Therefore, demagnetization intervals of 10-40 and 50-100 mT are chosen to 259 evaluate RPI signals recorded by the detrital and biogenic components, respectively. Due to the 260 logarithmic spacing of demagnetization steps, the 10-40 mT and 50-100 mT demagnetization 261 intervals are represented by the 10.7-41.6 mT and 52.8-100 mT measurements, respectively. The 262 slopes of NRM/ARM derived for the detrital and biogenic components have similar RPI variations 263 (Figure 6a), which indicates that a robust RPI signal is likely to have been recorded in core NP02 264 sediments, despite the presence of two magnetite sub-populations. 265



Figure 6. RPI estimates in core NP02. (a) RPI estimates from the detrital (NRM<sub>10-40 mT</sub>/ARM<sub>10-40 mT</sub>, blue curve) and biogenic (NRM<sub>50-100 mT</sub>/ARM<sub>50-100 mT</sub>, red curve) components with 95% confidence interval, respectively (filtered with a 15 cm cutoff); (b) cross-correlation results for the detrital and biogenic RPI estimates; and (c) distribution of relative offsets between the positions of peaks and troughs in detrital and biogenic RPI estimates.

## **3.4 Major element composition of glass shards**

Three visible tephra layers are identified at depths of 137-143 cm, 220-227 cm, and 391-395 cm. Each tephra layer has homogeneous glass populations with relatively narrow compositional ranges (Figure 7), which excludes the possibility of tephra mixing due to bioturbation and/or simultaneous deposition. Major element compositions of the 34 analyzed glass shards fall into the rhyolite field (Figure 7a). In terms of potassium contents, tephra layers 1 (137-143 cm) and 3 (391-395 cm) belong to the low-K field while tephra layer 2 (220-227 cm) falls into the medium-K field (Figure 7b).



**Figure 7.** Geochemical characteristics of the glass shards from core NP02. (a) Classification of the total alkali-silica (TAS) diagram with composition fields from Le Bas et al. (1986); and (b)  $SiO_2$ -K<sub>2</sub>O diagram with boundaries of high-, moderate, and low-K rocks according to Gill (1981). Error bars (2 $\sigma$ ) for the glass shards were calculated from the secondary glass standard (AHO-G).

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## 284 4 Discussion

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## 4.1 Age model construction

Age models are often constructed using interpolation/regression techniques, which assume 286 symmetrically/normally distributed age errors (Blaauw, 2010). However, <sup>14</sup>C calibrations often 287 result in asymmetrical and multi-peaked calendar age uncertainties (Blaauw, 2010; Buck et al., 288 1999; Ramsey, 1995), and such asymmetrical distributions can lead to problems for age model 289 construction. Consideration of full probability density functions can avoid such issues when 290 291 estimating age-depth models (Blaauw & Christen, 2005; Lougheed & Obrochta, 2019; Ramsey, 2008). There is also an inherent depth uncertainty in sampling geological archives. For core NP02, 292 an age model was constructed following Lougheed and Obrochta (2019) using age constraints 293 provided by the tephra correlation and one calibrated <sup>14</sup>C age in which probability density 294 functions are used, with incorporation of realistic errors for both age and depth. 295

Tephra are a product of large and explosive volcanic eruptions, which spread widely over 296 land, sea, and ice. Unless a volcanic ash has been reworked long after deposition, the same tephra 297 layer in different depositional sinks has effectively an identical age — an isochron — to within 298 about a year and is considered to be geologically synchronous (Lowe, 2011). Therefore, tephra 299 layers are excellent stratigraphic markers and are used widely for correlating and dating 300 sedimentary successions (e.g., Grönvold et al., 1995; Mangerud et al., 1984; Matsu'ura et al., 2014; 301 Rutledal et al., 2020; Smith et al., 2013). In the western North Pacific Ocean, tephra layers are 302 generally dominated by vitric material from high-SiO<sub>2</sub> magmas that originated in subduction-303 related volcanic events from the Japanese archipelago or the Kamchatka Peninsula (Horn et al., 304 1969; Ninkovich et al., 1966). By comparing glass shard chemistry results with data reported in 305 306 previous studies, each tephra layer from core NP02 can be correlated with known Japanese sourced tephra markers. Although Na<sub>2</sub>O contents do not overlap perfectly, major element compositions of 307 14 analyzed glass shards for tephra layer 1 are in good agreement with that of the To-Of tephra, 308 309 which has been dated to occur in late MIS 3 (32-43.2 ka) (Machida & Arai, 2003; Matsu'ura et al., 2014; Smith et al., 2013). The glass shard chemistry of tephra layers 2 and 3 are highly consistent 310 with that of the Spfa-1 and Kt-3 tephra, respectively (Table 1). The Spfa-1 tephra has an AMS <sup>14</sup>C 311 age of 45.105 - 46.560 cal. ka BP (Uesawa et al., 2016), while Kt-3 has an uncalibrated AMS  $^{14}$ C 312 age of  $54.616 \pm 0.857$  ka BP (Amma-Miyasaka et al., 2020). These tephra markers have been 313 314 reported widely in the North Pacific Ocean; co-occurrences of To-Of and Spfa-1 have been reported in cores MR99-K04 and MR98-03 (Aoki, 2000), which are close to core NP02 (Figure 315 316 1).

### 317 **Table 1.**

318 Average Major-element Compositions of Glass Shards from Core NP02 and Correlative Tephra Markers

319 (Normalized).

	Tephra 1		Tephra 2		Tephra 3		To-Of <sup>a</sup>		Spfa-1 <sup>a</sup>		Kt-3 <sup>a</sup>	
	Avg.	$\pm 1\sigma$	Avg.	$\pm 1\sigma$	Avg.	$\pm 1\sigma$	Avg.	$\pm 1\sigma$	Avg.	$\pm 1\sigma$	Avg.	±lσ
SiO <sub>2</sub>	74.52	2.16	77.74	0.40	75.14	0.48	77.1	0.74	77.36	0.75	74.73	0.45
TiO <sub>2</sub>	0.37	0.07	0.14	0.01	0.35	0.02	0.42	0.07	0.15	0.04	0.35	0.03
$Al_2O_3$	13.83	1.20	12.31	0.24	12.84	0.22	12.65	0.29	12.37	0.48	12.92	0.39
FeO	1.92	0.32	1.48	0.09	2.84	0.04	1.97	0.23	1.47	0.16	2.76	0.16
MnO	0.08	0.08	0.05	0.07	0.08	0.07	0.08	0.06	0.06	0.02	0.10	0.02
MgO	0.51	0.14	0.16	0.03	0.44	0.03	0.40	0.07	0.18	0.06	0.46	0.05
CaO	2.66	0.69	1.45	0.03	2.68	0.04	2.22	0.21	1.57	0.22	2.69	0.21
Na <sub>2</sub> O	4.76	0.27	4.20	0.15	4.19	0.18	3.91	0.12	4.10	0.19	4.37	0.17
$K_2O$	1.12	0.15	2.44	0.07	1.38	0.02	1.25	0.06	2.70	0.12	1.56	0.12
$P_2O_5$	0.05	0.05	0.03	0.02	0.06	0.03			0.03	0.03	0.06	0.03
Total	100		100		100		100		100		100	
n	14		15		5		15		114		30	

<sup>a</sup>Average compositions of To-Of tephra is from Machida and Arai (2003), Spfa-1 and Kt-3 tephras are from

321 Amma-Miyasaka et al. (2019).

The To-Of tephra does not provide a suitable age constraint for core NP02 because of its 322 large age uncertainty. However, its occurrence suggests that sediments above this layer were 323 deposited after 40 ka, a time when radiocarbon can play a role in chronology development. 324 Therefore, discrete samples (3 cm in size) were taken from the upper 137 cm of core NP02, and 325 were then washed, sieved, and observed under the microscope to seek foraminifera. Foraminifera 326 are generally preserved poorly in sediments deposited below the CCD, and foraminifera (N. 327 328 pachyderma sinistral) were only identified at a depth of 49-52 cm. A sample from this depth yielded a radiocarbon date of 28.413 (28.077-28.714,  $2\sigma$ ) cal. ka BP. 329

Ages for the upper (above 50 cm) and lower (below 395 cm) parts of core NP02 are estimated by polynomial fitting and linear regression, respectively. Based on upward extrapolation, the age estimate for the core-top (0 cm) is 21.07 ka. An absence of surficial sediments is common for gravity coring where the surface layer is removed by impact of the core on the seafloor (Adelseck & Anderson, 1978; Karlin, 1990; Skinner & McCave, 2003). After removal of 17 cm for the three tephra-containing sediment intervals, which are considered as instantaneous deposits, the mean sedimentation rate for core NP02 based on our age model is ~11.6 cm/kyr (Figure 8).



Figure 8. Age-depth model for core NP02. Ages for the upper (<50 cm), intermediate (50-395 cm), and lower (>395 cm) parts are estimated from polynomial fitting, probability density functions (Lougheed and Obrochta, 2019), and linear regression, respectively. Grey shaded areas and blue lines indicate the  $1\sigma$  and  $2\sigma$  confidence intervals, respectively. The radiocarbon date and tephra layers with age errors are also indicated.

342

#### 4.2 Core NP02 RPI characteristics and recording fidelity

Remanence lock-in in sediments can lead to depth offsets between the sediment/water interface and the recorded remanence (Tauxe, 1993; Verosub, 1977). Recent analyses of remanences carried by detrital and biogenic components indicate that there is no relative lock-in

depth offset between RPI signals recorded by these components within marine sediments (Chen et 346 al., 2017; Ouyang et al., 2014). For core NP02, cross-correlation results for the detrital and 347 348 biogenic components suggest no phase lag for the whole sediment sequence (Figure 6b). When comparing the depth positions of all RPI maxima and minima in the sequence, the distribution of 349 relative lock-in depth offsets is centered on zero offset (Figure 6c), which is consistent with the 350 cross-correlation results. However, relative lock-in variations produce a maximum depth error of 351 5 cm (Figure 6c). For continuous high-resolution RPI studies, any relative lock-in depth offsets 352 could be a crucial limitation when assessing leads and lags between millennial scale climatic 353 variations and, therefore, should be considered as part of the age uncertainty (Roberts et al., 2013). 354 The relative depth offset induces an age error <0.5 kyr for points in the RPI record from core NP02, 355 356 which should have little effect on comparisons with other RPI stacks/records. Moreover, there is no significant coherence between the RPI estimates and the normalizers, which suggests that 357 sedimentary and/or environmental modulation of the RPI signals were negligible (Figure S2). 358

Core NP02 RPI estimates are plotted with respect to age in Figure 9a. The RPI minima 359 360 centered at ~27 ka and ~43 ka in core NP02 are accompanied by aberrant inclinations (Figure S3), which are likely associated with the Rockall and Laschamp excursions (Bonhommet & Babkine, 361 1967; Bonhommet & Zähringer, 1969; Channell, Harrison, et al., 2016). The Laschamp excursion 362 363 is the most pervasively documented and studied Quaternary geomagnetic excursion, which occurred at ~41 ka and lasted less than 3 kyr (Laj et al., 2000; Wagner et al., 2000). During the 364 Laschamp excursion, inclination directional changes have been observed from locations with a 365 broad geographical coverage (e.g., Caricchi et al., 2019; Ferk & Leonhardt, 2009; Laj et al., 2006 366 and references therein; Nowaczyk et al., 2012). However, geomagnetic field behaviour during the 367 Laschamp excursion over the North Pacific Ocean seems to have been different from that over the 368

rest of the world; reversed polarity inclinations have not been recorded in NP02 (Figure S3) and 369 other North Pacific Ocean cores (e.g., Inoue & Yamazaki, 2010; Okada et al., 2005; Roberts et al., 370 1997; Shin et al., 2019; Xiao et al., 2020; Zhong et al., 2020). Also, a model for the Laschamp 371 excursion predicts that the field intensity over the Pacific Ocean remained relatively high 372 compared to most locations on Earth (Leonhardt et al., 2009). The different geomagnetic field 373 behaviour over the North Pacific Ocean during the Laschamp excursion may have resulted from a 374 relatively large non-dipole to dipole field ratio, which remains to be explored with the aid of 375 higher-resolution records. 376

Core NP02 RPI variations are compared with global RPI stacks (Sint-2000, PISO-1500, 377 and GLOPIS-75) and regional stacks from the equatorial (West Caroline Basin stack), mid-latitude 378 379 (the NOPAPIS-250 stack), and subarctic Pacific Ocean (Channell et al., 2009; Laj et al., 2004; Valet et al., 2005; Yamamoto et al., 2007; Yamazaki et al., 2008; Zhong et al., 2020) (Figure 9a). 380 381 The RPI minimum centered at ~43 ka in core NP02 correlates with the marked RPI Laschamp excursion minimum at ~40-42 ka identified in all other stacks. The lag in recording of this RPI 382 383 minimum in core NP02 may be due to age model imprecision and/or variable sedimentation rates. 384 Before the Laschamp excursion, all records have broad RPI maxima in the  $\sim$ 47-55 ka interval. After the Laschamp excursion, the records have a generally increasing trend to ~30 ka. Although 385 386 the reference stacks generally suggest a similar overall pattern, there are still differences in their fine detail. For example, two peaks at ~47 ka and 53 ka occur in the GLOPIS-75 stack, which are 387 not observed in the other stacks that have a broad maximum at ~47-55 ka. Also, the subarctic 388 Pacific stack has larger variations after the Laschamp excursion than the other stacks, particularly 389



390 Figure 9. Comparisons of core NP02 RPI estimates with (a) global stacks (Sint-2000 (Valet et al., 2005), PISO-1500 (Channell et al., 2009), and GLOPIS (Laj et al., 2004)); regional stacks from the equatorial 391 (West Caroline Basin stack, Yamazaki et al., 2008), mid-latitude (NOPAPIS-250, Yamamoto et al., 2007), 392 and subarctic Pacific Ocean (Zhong et al., 2020); and (b) individual RPI records of core BOW-8A (Okada 393 394 et al., 2005), Site 883/884 (Roberts et al., 1997; Roberts, 2008), core GC1+PC7 (Inoue & Yamazaki, 2010; Yamazaki et al., 2013), core NGC65 (Yamazaki, 1999), and core NPGP1401-2A (Shin et al., 2019). The 395 Laschamp excursion is denoted by the pink shaded area. Positions of the radiocarbon date (grey triangle) 396 397 and tephra layers (brown triangle) from core NP02 are indicated in (a). All stacks and RPI records are 398 rescaled to the 0-1 range. Dashed lines in (b) denote possible correlations between core NP02 and GC1+PC7.

compared to the other regional Pacific stacks. Zhong et al. (2020) attributed the contrasting 399 subarctic North Pacific RPI pattern to regionally variable geomagnetic field behaviour. However, 400 this large variability is not observed in the geomagnetic model of Ziegler and Constable (2015), 401 which covers a longer time duration and suggests lower variability and less significant recording 402 of excursions in the Pacific Ocean compared to the Atlantic Ocean (Ziegler & Constable, 2015). 403 Given that stacking of multiple records with variable resolution can distort amplitude features and 404 smooth high-frequency and local geomagnetic variations, we further compare individual RPI 405 records from the North Pacific Ocean and its marginal seas (Inoue & Yamazaki, 2010; Okada et 406 al., 2005; Roberts, 2008; Roberts et al., 1997; Shin et al., 2019; Yamazaki, 1999; Yamazaki, Inoue, 407 et al., 2013) (Figure 9b). To aid comparison, only records with geomagnetism-independent age 408 409 models are considered. The RPI record of core NP02 resembles that of GC1+PC7, which validates the core NP02 RPI signal. Apart from this comparison, consistency among individual records is 410 less obvious than among the stacks, which may be associated with age model imprecisions and 411 412 variable sedimentation rates. Nevertheless, pronounced RPI amplitude variations are recorded at Site 883/884, core GC1+PC7, and core NP02 from mid to high latitudes. In contrast, core BOW-413 414 8A from the Bering Sea and cores from latitudes lower than 40° N have less significant amplitude 415 variations, and it is difficult to ascertain whether the lower variability is a result of genuine 416 geomagnetic field behaviour or imperfect recording that smooths the large variability due to low sedimentation rates. The model of Ziegler and Constable (2015) is based mainly on records from 417 the equatorial Pacific Ocean, so their documented low field variability over the Pacific Ocean 418 419 could be biased; more records from the mid to high latitude Pacific Ocean are required to improve our understanding of Pacific geomagnetic field behaviour. 420

### 421 **5** Conclusions

Dating of Quaternary North Pacific Ocean sediments is challenging because they were 422 usually deposited below the CCD. A geomagnetism-independent age model for North Pacific 423 Ocean core NP02 is established with the aid of tephra correlation (using the To-Of, Spfa-1, and 424 Kt-3 tephra layers) and one radiocarbon date. Detrital vortex state and biogenic SD magnetite in 425 sediments from core NP02 both record RPI variations faithfully, which indicates that RPI-assisted 426 chronology is a promising dating tool for North Pacific Ocean sediments. To assess the reliability 427 428 of RPI determinations, apart from the common criteria suggested by Tauxe (1993), assessment of component-specific magnetic properties, as presented here, is also required. For sediments that 429 contain more than one magnetic mineral component, the RPI record should be checked to ensure 430 431 that each component records consistent RPI variations. Moreover, relative depth offsets in recording of components needs to be checked to assess whether this provides further age 432 433 uncertainties that should be taken into consideration.

434

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### 447 Data availability statement

448 Data produced in this study are available at Mendeley Data
449 (<u>https://data.mendeley.com/datasets/42pyc27gy7/1</u>) via DOI:10.17632/42pyc27gy7.1.
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# Figure 1



Figure 2





Figure 4









Figure 8



# Figure 9

